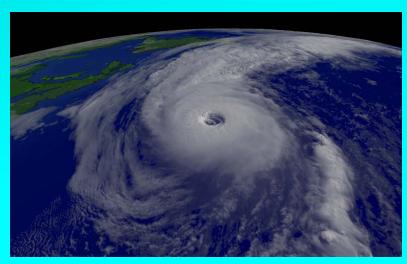
IMAGE DEBLURRING - COMPUTATION OF CONFIDENCE INTERVALS

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Introduction



Andromeda nebulae, picture from NASA



Hurricane Alex, picture from NOAA

Blurring

- Machine errors in transforming the image into data
- Background (Light , contrast)
- Motion of the object, motion of the camera

Example 1



Clear Image

Blurred Image

Example 2



Clear Image

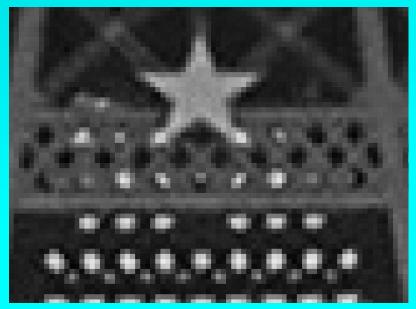
Blurred Image



Pixels

Pixel Values [0,255]

0: black 255: white





Notation

Symbol	Size	Explanation
K	mxn	Matrix defined through the Point Spread function (PSF) in the case of a linear problem
Х		Original Clear Image
X	nx1	Vector containing the values corresponding to the pixels of the image X
В		The blurred image we measure
b	mx1	Vector which contains the values of the pixels of the blurred image B
е	mx1	Vector of noise

Model

In the case of linear Point Spread Function

$\mathbf{b} = \mathbf{K}\mathbf{x} + \mathbf{e}$

Problem

• Given **b**, matrix **K** and a distribution for the noise **e** with mean value 0 and variance a matrix S^2 , we need to compute confidence intervals for the quantities $\varphi_k^* = \mathbf{w}_k^T \cdot \mathbf{x}$, for k = 1, 2, ..., p where **w**_k are given vectors.

Confidence intervals

- Intervals into which the values of a parameter fall, with a certain probability.
- Example: $100\alpha\%$ confidence interval, $\alpha \le 1$ One at a time confidence intervals $\Pr\{l_k \le \varphi_k^* \le u_k\} = \alpha, k = 1, 2, ..., p$

Simultaneous confidence intervals

$$\Pr\{l_k \le \varphi_k^* \le u_k , k=1,2,\ldots,p\} \ge \alpha$$

• Suppose that the noise is normally distributed, and that $\hat{\mathbf{x}}$ is an unbiased estimate of the true solution . Then, given α in (0,1), there is a 100 α % probability that the true value of $\mathbf{W}^T \mathbf{x}$ is contained in the interval [l, u]where $l = \min_{\mathbf{x}} \{ \mathbf{w}^T \mathbf{x} : \| \mathbf{K} (\mathbf{x} - \hat{\mathbf{x}}) \|_{\mathbf{S}}^2 = \kappa^2 \}$ and , $u = \max_{\mathbf{x}} \{ \mathbf{w}^T \mathbf{x} : \| \mathbf{K} (\mathbf{x} - \hat{\mathbf{x}}) \|_{\mathbf{S}}^2 = \kappa^2 \}$

Where $\|\mathbf{K}(\mathbf{x}-\hat{\mathbf{x}})\|_{s}^{2} = [\mathbf{K}(\mathbf{x}-\hat{\mathbf{x}})]^{T} \mathbf{S}^{-2} [\mathbf{K}(\mathbf{x}-\hat{\mathbf{x}})]$ and κ is such that for the probability α , i.e., for confidence $100\alpha\%$ $\alpha = \int_{-\kappa}^{\kappa} n(\mathbf{x};0,1)d\mathbf{x}$ where $n(\mathbf{x};0,1)$ is the normal

distribution of \mathbf{x} with mean value o and variance 1.

• With the same assumptions, $\mathbf{x} \ge 0$ and \mathbf{S} nonsingular and symmetric, the probability that φ^* is contained in the interval [l, u] is greater than or equal to α where $l = \min \left\{ \mathbf{w}^T \mathbf{x} : \left\| \mathbf{K} (\mathbf{x} - \hat{\mathbf{x}}) \right\|_{\mathbf{S}} \le \mu, \mathbf{x} \ge 0 \right\}$ and $u = \max \left\{ \mathbf{w}^T \mathbf{x} : \left\| \mathbf{K} (\mathbf{x} - \hat{\mathbf{x}}) \right\|_{\mathbf{s}} \le \mu, \mathbf{x} \ge 0 \right\}$ $\operatorname{rank}(\mathbf{K}) = q \, \alpha = \int \chi_q^2(\rho) d\rho \, r_0 = \min_{\mathbf{x}} \left\| \mathbf{K}(\mathbf{x} - \hat{\mathbf{x}}) \right\|_{\mathbf{S}}^2 \, \mu^2 = r_0 + \gamma^2$ and χ_a^2 is the probability density function for the chi-squared distribution with q degrees of freedom.

The matrix **K**

- The matrix **K** is defined through the Point Spread function (PSF) and can be determined by using a single point source in a pixel of an image and then moving it to other pixels of the image to obtain more blurred images. Knowing the clear image and the blurred image for all the pixels we can compute the matrix **K**.
- In the following, the matrix **K** will be a given nxn square matrix.

Sub-matrices

We want to reduce the size of the matrix (and so of the image) to facilitate the computations. Suppose that b = Kx.
 For a sub-image rxc we proceed as follows: We have two matrices E and E where E has rc columns of unit vectors and E has n-rc columns of unit vectors. Then,

$$\mathbf{E}^{T}\mathbf{K}\mathbf{x} = \mathbf{E}^{T}\left(\mathbf{K}[\mathbf{E}\overline{\mathbf{E}}]\right)\left(\begin{bmatrix}\mathbf{E}^{T}\\\mathbf{E}^{T}\end{bmatrix}\mathbf{x}\right) = \left(\mathbf{K}_{s} \quad \mathbf{K}_{t}\right)\left(\begin{bmatrix}\mathbf{X}_{s}\\\mathbf{X}_{t}\end{bmatrix}\right) = \mathbf{K}_{s}\mathbf{X}_{s} + \mathbf{K}_{t}\mathbf{X}_{t}$$

where \mathbf{X}_{s} is the vector corresponding to the sub-image of the original image. Ignoring the second term of the right hand side, we obtain $\mathbf{b}_{s} = \mathbf{K}_{s} \mathbf{x}_{s}$

Implementation

- Matlab (IPT- Image Processing toolbox)
- Case of spatially invariant blur
 Smaller Problems- same matrix
 Parallel Computing (Open MP or MPI)

Databases

<u>http://sipi.usc.edu/database/</u>
 USC-SIPI Image Database
 Signal & Image Processing
 Institute of the University
 of Southern California



 <u>http://www.imageprocessingplace.com/root_files_V3/</u> <u>image_databases.htm</u>
 Image Processing Place

Validation

- Known data -----> Expected results
- Take images
- Blur images
- Add noise
- Use code to compute the confidence intervals for these images
- Deblur images
- Count samples in the computed intervals
- If close to 95% fall in the 95% confidence intervals, then the code is validated.
- If not... check code, correspondence with theory and special circumstances

Testing

Good code: Right results in relatively short time

Depends on

- Size of image
- Format of image
- Various Point Spread Functions

Schedule-Milestones

- September: literature, image processing toolbox of Matlab, Project Proposal
- October: literature, problem, code
- November: code, validation
- Early December: midyear report
- Late January: testing
- February: parallel computing
- March: testing, validation
- April: validation, final report
- May: final presentation

REFERENCES

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Thank you